



## **PETROGENESIS OF HOMRAT EL-GIRIGAB ALKALI-FELDSPAR GRANITES, NORTHERN EASTERN DESERT, EGYPT**

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### **ABSTRACT**

Homrat El-Girigab area located at Northern Eastern Desert which, characterized by abundant intrusion of calc-alkaline and alkaline/peralkaline granitoids and their associated volcanics. These granitoids have a particular geodynamic interest as they provide an outstanding opportunity to tell how continental crust of ANS was formed. Homrat El-Girigab area is covered by Dokhan volcanics (andesite & dacites), which intruded by alkali-feldspar granites. The chemistry of biotites indicates that, the alkali-feldspar granites were crystallized from alkaline crustal source under oxidized conditions (i.e. nickel-nickel oxide buffer or NNO). They were crystallized under conditions including, temperatures range from 700 to 750 °C, pressures 3 to 4 kbar, depths of emplacement range from 7 to 11 km and under Oxygen fugacity ( $\log fO_2$ ) ranges from -15 to -16. Homrat El-Girigab alkali-feldspar granites (HGAFGs) are alkaline, ferroan anorogenic (i.e. extensional) A-type granites. They were emplaced during the late post-collisional crustal extensional stage at which the effect of lithospheric delamination, and thus asthenospheric uprise, likely diminishes. At this stage the mantle-derived mafic melts start intraplating the lower crustal levels, that facilitated by the abundance of strike-slip faults and shear zones. This lithospheric intraplating caused widespread melting producing the alkaline magma of HGAFGs. The studied granites were derived from lower crustal amphibolitic source and evolved mainly by fractional crystallization.

**Keywords:** Homrat El-Girigab, Eastern Desert, Mineral chemistry, Petrogenesis

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### **INTRODUCTION**

The Arabian Nubian Shield (ANS), represents the biggest Neoproterozoic juvenile continental crust belt on Earth. The evolution ANS juvenile crust includes three subsequent stages: 1) subduction (~870-670 Ma), including the formation of island arc volcano-sedimentary sequences, plutonic rocks and amalgamation of these accreted terrains onto East Gondwana continental block (Abdelsalam & Stern, 1996; Condie et al., 2009); 2) continental collisional (650-640 Ma) between the juvenile accreted ANS crust with pre-Neoproterozoic continental blocks of West Gondwana (Abdelsalam et al., 2002), which characterized by formation of calc-alkaline gabbros and granodiorites; 3) late to post-collisional (590–550 Ma), which follow collision involving extensional collapse of the thickened lithosphere (Avigad et al., 2005; El-Bialy, 2010; Eyal et al., 2010; Be'eri-Shlevin et al., 2011), including formation of volcanics and calc-alkaline to alkaline/peralkaline granitic rocks.

The Egyptian granitoids were classified into: 1) older, syn-orogenic calc-alkaline, I-type granitoids (880-610 Ma), which related to the end of Pan-African orogeny; 2) younger post-orogenic either not strictly anorogenic, calc-alkaline/alkaline granitoids (640-610 Ma) or alkaline/peralkaline, true anorogenic A-type granites (600-550 Ma). The later are related to rifting (Bielski et al., 1979; Moghazi, 1999, Azer, 2013). Different models have been proposed for the sources of the post-collisional A-type granites in ANS, which invoke contrasting continental crust and mantle reservoirs (Ali et al., 2009, 2014; Be'eri-Shlevin et al., 2009, 2010; El-Bialy & Hassen, 2012; Eyal et al., 2010; Farahat & Azer, 2011; Moghazi et al., 2011, 2012).

Homrat El-Girigab is located at northern Eastern Desert terrain, which characterized by abundant intrusion of calc-alkaline and alkaline/peralkaline granitoids and their associated volcanics, constituting

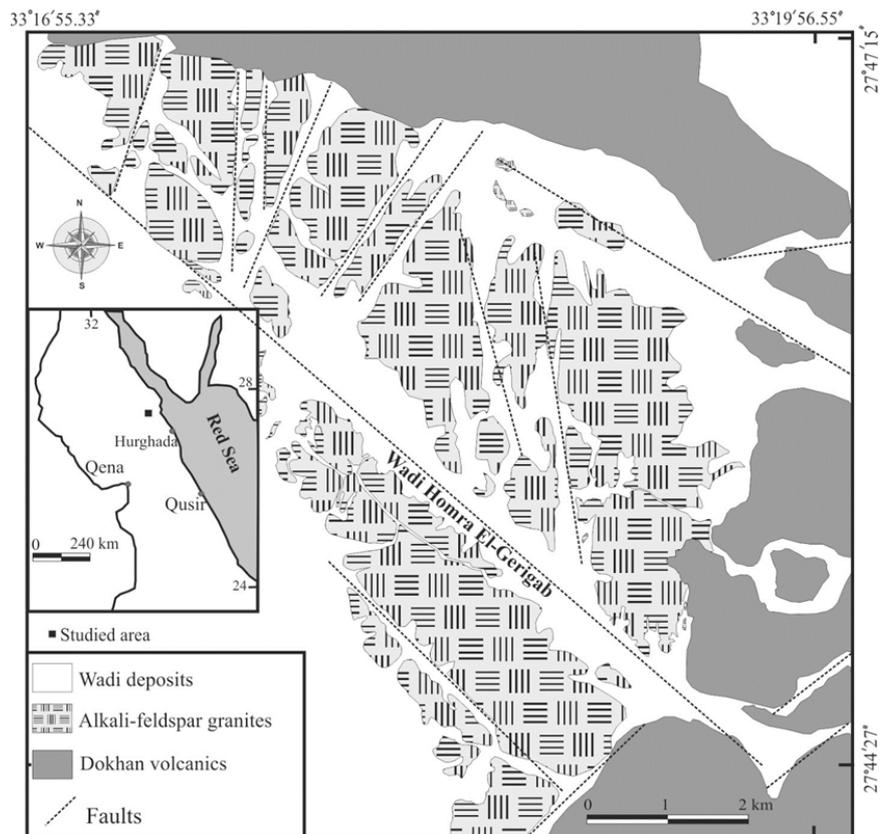
~80% of the basement outcrops, therefore, these granitoids have a particular geodynamic interest as they provides an outstanding opportunity to tell how continental crust forms. In addition, some of late to post-collisional A-type granite plutons in ANS are considered specialized granites (Drysdall et al., 1984; Küster, 2009; Johnson et al., 2011). A-type granites in ANS received much attention due the potential economic significance of their elevated abundances of REE, U, Th, Ta and Nb. Homrat El-Girigab area got more attention due to the high radioactivity of both its uraniferous granite and the stream sediment along the main Wadi (Salman et al., 1995; Abd El-Ghafour et al., 2001). Furthermore, the granitic rocks in the area are mined for feldspar at latitude 27° 45' 30" and longitude 33° 18' 20" as raw material for ceramic industry. Wetait et al. (2014) described the granitic rocks in Homrat El-Girigab as calc-alkaline I-type granites, which were derived in a post-orogenic tectonic environment.

The aim of the present work is to detect the petrogenesis of the Homrat El-Girigab alkali-feldspar granites through their mineralogical and geochemical characteristics.

### GEOLOGIC SETTING

Homrat El-Girigab area located in the northern part of Esh El-Mallaha range, Northern Eastern Desert. The area lies about 70 km to southeast of Ras Ghareb city, 15 km to southwest of Cairo- Hurghada Road. It is located between latitude 27° 44' 27"- 27° 47' 15" N and longitude 33° 16' 55.33" - 33° 19' 56.55" E. Homrat El-Girigab area is covered by moderate relief peaks of Dokhan volcanics (DVs), which are intruded by high relief hills of alkali-feldspar granites (Fig. 1& Fig. 2a). Dokhan volcanics are hard, massive and non-foliated rocks. They show grey, greenish grey to dark black colors and they are represented by andesite and dacites varieties.

Fig.1: Geological map of Homrat El-Girigab area.



HGAFGs peaks are surrounded by Dokhan volcanic hills. They are hard massive with reddish pink to red color, locally weathered, well jointed and show exfoliation joints (Fig. 2b). HGAFGs are dissected by acidic dykes (i.e. felsite and porphyritic rhyolite). These granitic rocks contain sub-rounded xenoliths of Dokhan volcanics. The area is dissected by major faults trending NW-SE and NE-SW and subordinate NNW-SSE and E-W trends.

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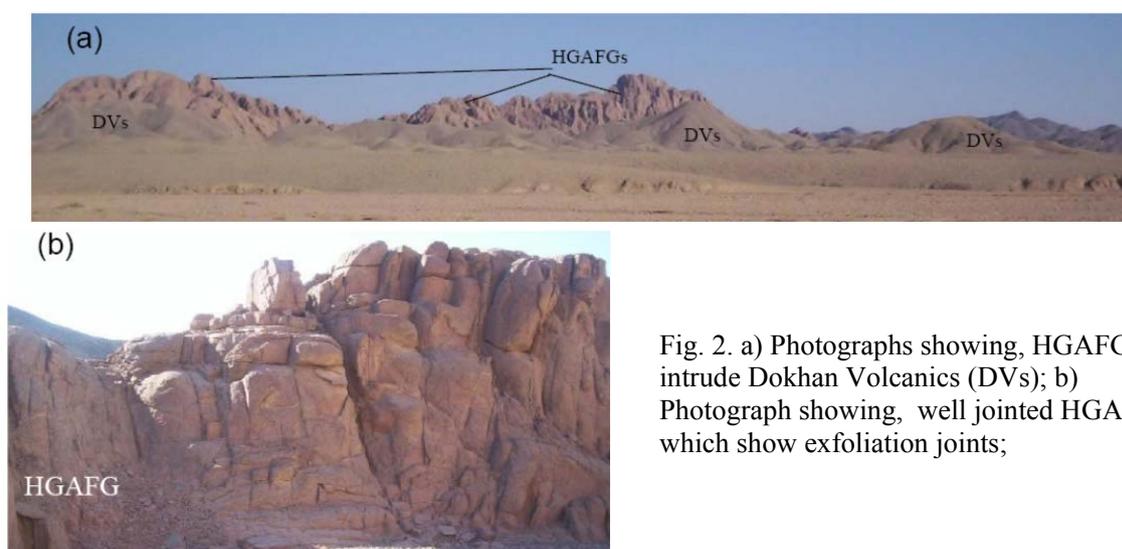


Fig. 2. a) Photographs showing, HGAFGs intrude Dokhan Volcanics (DVs); b) Photograph showing, well jointed HGAFGs, which show exfoliation joints;

### PETROGRAPHY

HGAFGs are medium to coarse-grained hypidiomorphic rocks, being composed mainly of K-feldspar, quartz, plagioclase with subordinate amount of biotite. The accessory minerals are sphene, zircon, apatite, allanite and fluorite, while the secondary minerals are kaolinite, sericite and chlorite. The rock shows porphyritic, rapakivi and granophyric textures.

K-feldspars are represented by subhedral crystals of orthoclase-perthite and microcline. Some perthite crystals are mantled by oligoclase forming rapakivi texture (Fig. 3a). Quartz occurs as subhedral to anhedral crystals, which invade with reaction rim along their contacts. Granophyric intergrowth between K-feldspar and quartz is common (Fig. 3b). Plagioclase occurs either as medium to coarse subhedral tabular crystals. Plagioclase crystals are corroded and enclosed optically by other mineral constituents. Biotite occurs as subhedral to anhedral flacks (Fig. 3c), which are partly altered to chlorite. Sphene occurs as subhedral crystals filling the interstices between quartz and feldspar (Fig. 3d). Zircon occurs as subhedral to euhedral crystals which associated with apatite and allanite (Fig. 3e). Apatite form colorless euhedral prismatic crystals. Allanite occurs as subhedral crystals which are sending anastomosing cracks along its surrounding minerals. Fluorite occurs as subhedral to anhedral grains associated with biotite (Fig. 3e).

### Analytical techniques

A total of 7 spots in biotite from alkali-feldspar granites were analyzed by microprobe analyses (Table 1). The microprobe analyses were carried out using Cameca SX-100 Electron probe Microanalyzer, at Department of Electron Microanalysis, Geological Institute of Dionyz Stür, Bratislava, Slovak Republic. The microprobe was operated at an acceleration voltage of 15 kv, beam current of 20 nA, electron beam diameter of 5  $\mu\text{m}$  and using ZAF corrections. The chemical analyses of 6 fresh representative samples from HGAFGs (Table 2) were carried out using Panalytical Axios Advanced X-ray fluorescence (XRF). The analytical precision is 2-5%. The chemical analyses were carried out at Central Metallurgical Research and Development Institute (CMRDI), El-Tebbin, Helwan, Egypt.

### Mineral chemistry

The analyzed biotites show a limited variation in composition and comparable contents of  $\text{Al}_2\text{O}_3$ ,  $\text{TiO}_2$ ,  $\text{FeO}$  and  $\text{MgO}$  (Table 1). Fluorine content ranges from 0.44 to 1.86 (Table 1). According to  $\text{FeO} + \text{MnO} - 10*\text{TiO}_2 - \text{MgO}$  ternary diagram (Nachit et al., 2005) the studied trioctahedral micas are of primary magmatic origin (Fig. 4a). By using  $\text{Al}^{\text{I}}\text{-Fe}^{2+}/(\text{Fe}^{2+} + \text{Mg})$  binary diagram (Rieder et al., 1999), the studied micas were classified as biotite (Fig. 4b).

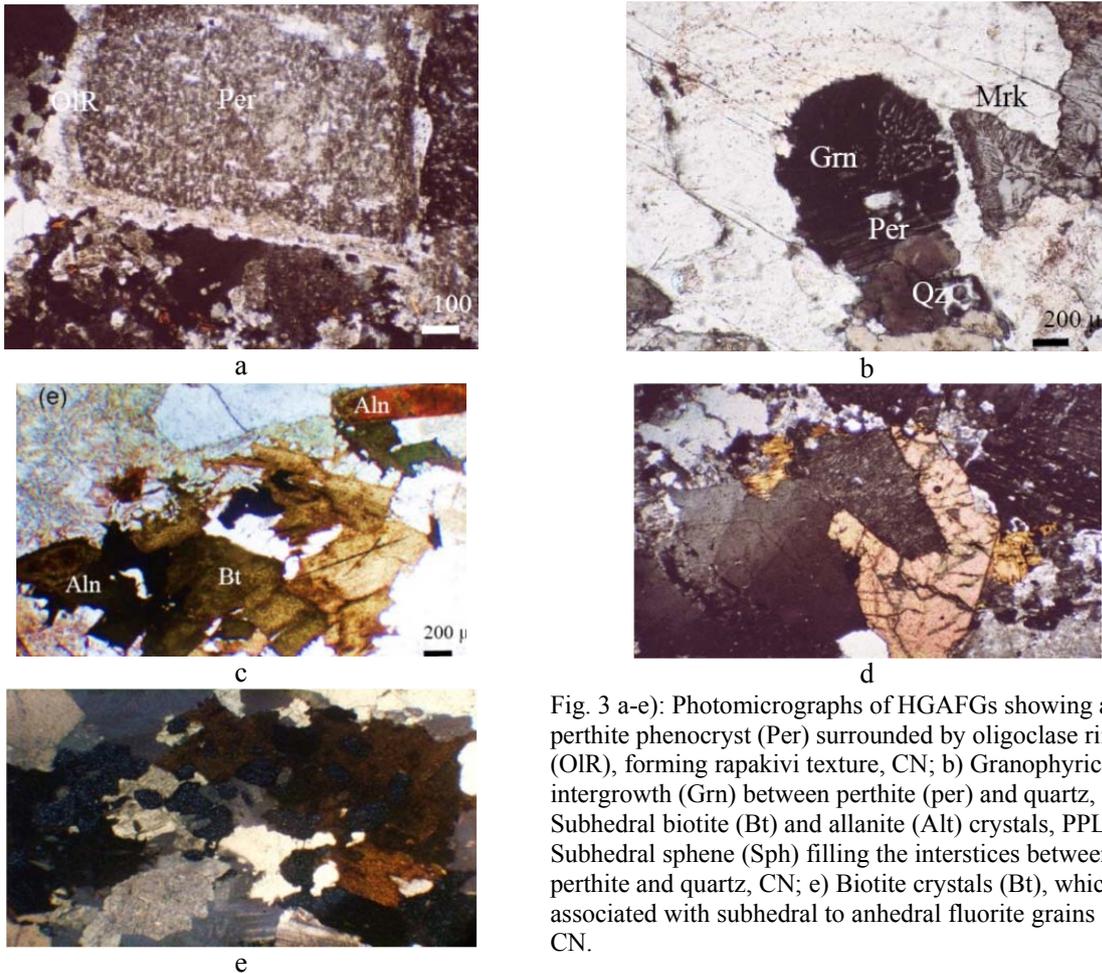


Fig. 3 a-e): Photomicrographs of HGAFGs showing a) perthite phenocryst (Per) surrounded by oligoclase rim (OIR), forming rapakivi texture, CN; b) Granophytic intergrowth (Grn) between perthite (per) and quartz, CN; c) Subhedral biotite (Bt) and allanite (Aln) crystals, PPL; d) Subhedral sphene (Sph) filling the interstices between perthite and quartz, CN; e) Biotite crystals (Bt), which associated with subhedral to anhedral fluorite grains (Flu), CN.

As aluminum play an important role in the alkalinity of the magma and could be used as a factor controlling alkalinity, acidity and as an indicator for pressure prevailing during the process of crystallization, thus the biotite composition could be used to detect the nature of magmatic source. On Mg- Al<sup>t</sup> binary diagram (Nachit et al., 1985), the biotite in HGAFGs implies crystallization from an alkaline magma (Fig. 3c).  $Fe^{2+}/(Fe^{2+} + Fe^{3+})$  ratios range from 0.86 to 0.90 (Table 1). Furthermore, the analyzed biotite samples were clustered around nickel-nickel oxide (NNO) buffer in  $Fe^{2+}$ -  $Fe^{3+}$ - Mg ternary diagram (Wones & Eugester, 1965), which suggest a moderate oxidation conditions (Fig. 4d). The calculated Oxygen fugacity buffers ( $Log fO_2$ ) of the studied biotite range from -14.92 to -16.28 according to the equilibrium equation of Huebner & Sato (1970) (Table 1). By using the calculated Oxygen fugacity buffer and the  $Fe^{2+}/(Fe^{2+} + Fe^{3+} * 100)$  ratio the biotite analyses were plotted on  $T^{\circ}C$ -  $Log fO_2$  diagram (Wones & Eugester, op. cit.). The studied biotites follow the oxidizing trend (Fig. 4e), which represent the undersaturated conditions during the crystallization of biotites from the melt (Wones & Eugester, 1965). The temperatures of crystallization that calculated from the chemistry of biotites range from 700 to 750.

## GEOCHEMISTRY

### Major and trace element

As shown in Table (2), HGAFGs have limited compositional variation, exhibiting narrow ranges of  $SiO_2$  (70.01- 72.91 wt. %),  $Al_2O_3$  (13.68- 14.25 wt. %), CaO (0.66- 0.96 wt. %) and  $TiO_2$  contents (0.11- 0.23 wt. %). They are rich in alkalis ( $Na_2O + K_2O$ ), (10.67-11.41 wt. %) and depleted in MgO (0.13-0.31 wt. %). The enrichment of  $Fe_2O_3^t$  relative to MgO indicates A-type signature (Frost et al., 2001). The presence of sodium metasilicates (NS = 1, on average; Table 2) in the CIPW normative values is a characteristic feature of peralkaline rocks.

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Table 1. Microprobe analyses (wt. %) of and structural formula (apfu) biotite in HGAFGs

Sample Spot No.	21			33		18	
	1	2	3	1	2	1	2
Oxides							
SiO <sub>2</sub>	36.13	35.72	35.52	35.77	35.77	35.83	35.51
TiO <sub>2</sub>	3.54	3.36	3.12	3.08	3.08	2.95	3.76
Al <sub>2</sub> O <sub>3</sub>	11.42	11.26	11.42	11.22	13.31	13.23	13.47
FeO	27.54	28.89	29.66	29.08	29.48	29.53	28.76
MgO	5.57	5.18	5.44	5.52	2.59	2.43	2.56
MnO	0.58	0.53	0.40	0.42	0.49	0.50	0.45
CaO	0.02	0.01	0.01	0.00	0.03	0.01	0.05
Na <sub>2</sub> O	0.03	0.03	0.13	0.10	0.11	0.10	0.16
K <sub>2</sub> O	8.79	8.99	8.44	8.61	8.92	8.92	8.84
F	1.13	1.10	0.97	0.44	1.29	1.86	1.52
Cl	0.37	0.42	0.41	0.42	0.47	0.47	0.52
H <sub>2</sub> O	1.45	1.51	1.48	1.68	1.46	1.24	1.39
Total	96.17	96.44	96.21	96.05	96.34	95.97	96.24
Structural Formula based on 22 Oxygen atoms							
Si	6.02	5.88	5.88	5.84	5.89	6.09	5.86
Al(IV)	1.98	2.12	2.12	2.16	2.11	1.91	2.14
Al(VI)	0.26	0.06	0.11	0.00	0.47	0.74	0.47
Ti	0.44	0.42	0.39	0.38	0.38	0.38	0.47
Fe <sup>3+</sup>	0.48	0.44	0.43	0.43	0.53	0.55	0.55
Mn	0.08	0.07	0.06	0.06	0.07	0.07	0.06
Mg	1.38	1.27	1.34	1.34	0.64	0.62	0.63
Fe <sup>2+</sup>	3.35	3.54	3.67	3.54	3.53	3.65	3.42
Na	0.01	0.01	0.04	0.03	0.03	0.03	0.05
K	1.87	1.89	1.78	1.79	1.87	1.93	1.86
Ca	0.00	0.00	0.00	0.00	0.01	0.00	0.01
Fe <sup>2+</sup> /Fe <sup>2+</sup> +Fe <sup>3+</sup>	0.87	0.89	0.90	0.89	0.87	0.87	0.86
T °C	750.56	737.33	729.59	727.20	700.41	700	728.69
Log (fO <sub>2</sub> )	-14.92	-15.09	-15.36	-15.51	-16.09	-16.28	-15.53

Homrat El-Girigab granites classified as alkali-feldspar granite (Fig. 5a) on ANOR-Q diagram (Streckeisen & Le Maitre, 1979). Eyal et al. (2010) distinguished two sub-stages of alkaline suites in Sinai: a) an early sub-stage (~610–600 Ma) formed syenogranites, which were preceded by monzodiorite; b) a later alkaline sub-stage (~608–580 Ma) including syenogranites, alkali-feldspar granites and peralkaline riebeckite granites. This later sub-stage was heralded by extensive volcanic activity (i.e. comparable with Dokhan volcanics in Homrat El-Girigab area). Furthermore, they suggested that, the alkaline granites were evolved mainly by fractional crystallization. The composition of the studied granites is compared Umm Shommer and Sharm alkali-feldspar granites (from Eyal et al., op. cit.). HGAFGs are so close in their composition to later sub-stage rocks (Table 2).

High Apatitic Index value (AI = 1.1) of HGAFGs, enhances their alkaline natures (Liégeois et al., 1998). This alkaline affinity is further confirmed on SiO<sub>2</sub> - Na<sub>2</sub>O + K<sub>2</sub>O - CaO diagram (Frost & Frost, 2008; Frost et al., 2001) and major elements discrimination diagram of Sylvester (1989), (Figs. 5b & c, respectively). They are typical ferroan A-type granites as deduced from Figure 5d (Frost et al., 2001). This conclusion is strengthened by applying the classification diagrams of Whalen et al. (1987) and Eby, (1992), where they plot consistently in the fields of A-type granites (Figs. 5e and 4f, respectively). Their plotting in Field A<sub>2</sub> (4f after Eby, 1992) implies their mantle-derived origin in an anorogenic tectonic setting.

H<sub>2</sub>O were determined by stoichiometry calculations. Fe<sup>3+</sup>/Fe<sup>2+</sup> were calculated based on the method of Lin and Peng (1994); (T °C) temperature of crystallization and Oxygen fugacity buffer Log (fO<sub>2</sub>) of biotites calculated according to the equation of Huebner & Sato (1970).

Table 2. Major elements (wt. %), trace elements (ppm) analyses of HGAFGs.

Rock	Alkali-feldspar granites						Average	US AFG	SH AFG
Sample No.	33	21	20	18	14	9			
Oxides									
SiO <sub>2</sub>	70.01	70.78	71.97	70.74	72.41	72.91	71.47	69.5	71.69
TiO <sub>2</sub>	0.23	0.20	0.11	0.17	0.13	0.14	0.16	0.55	0.30
Al <sub>2</sub> O <sub>3</sub>	13.73	14.25	13.79	13.68	13.85	13.79	13.85	14	13.41
Fe <sub>2</sub> O <sub>3t</sub>	1.78	1.71	0.81	1.44	0.93	1.30	1.33	1	1.69
MnO	0.04	0.03	0.01	0.03	0.02	0.03	0.03	0.06	0.08
MgO	0.31	0.24	0.16	0.24	0.18	0.13	0.21	0.2	0.29
CaO	0.96	0.60	0.75	0.68	0.66	0.80	0.74	0.6	0.83
Na <sub>2</sub> O	5.40	6.03	5.78	5.43	6.30	5.33	5.71	4.6	4.16
K <sub>2</sub> O	5.85	5.38	5.21	5.49	4.54	5.34	5.30	5.2	5.53
P <sub>2</sub> O <sub>5</sub>	0.05	0.03	0.01	0.01	0.01	0.01	0.02	-	0.06
L.O.I	1.10	0.30	0.80	1.70	0.60	0.60	0.85	0.3	0.61
Total	99.46	99.55	99.40	99.62	99.65	100.37	99.68	99.85	99.97
F	1310	1620	1720	1380	1680	1690	1566.67	-	-
Ga	50	30	30	40	40	45	39.17	16	25
Ba	240	120	102	94	35	105	116.00	332	317.5
Rb	300	280	330	410	200	250	295.00	135	137.5
Sr	90	60	50	67	30	25	53.67	57	43
Nb	70	90	110	140	110	100	103.33	15	42.5
Y	89	90	78	74	69	87	81.17	24	50
Zr	290	310	250	270	200	230	258.33	180	
Zn	80	60	30	30	40	40	46.67		
Th	10	30	50	20	20	15	24.17	12.7	15.25
Q	20.15	19.17	22.19	21.32	22.33	23.52	21.45		
Or	34.58	31.79	30.77	32.47	26.82	31.58	31.34		
Ab	38.01	43.34	41.92	39.77	45.99	41.17	41.70		
Ns	1.79	1.79	1.63	1.43	1.70	0.92	1.54		
Hy	0.76	0.59	0.40	0.60	0.46	0.32	0.52		
Il	0.08	0.06	0.02	0.06	0.04	0.07	0.06		
Ru	0.19	0.17	0.10	0.14	0.11	0.10	0.14		
Ap	0.12	0.08	0.02	0.03	0.03	0.01	0.05		
Fr	1.24	0.78	1.03	0.93	0.90	1.10	1.00		
mg #	25.42	21.50	28.18	24.76	28.20	16.34	24.07		
AI	1.11	1.10	1.10	1.09	1.10	1.06	1.09	0.94	0.92
T <sub>Zr</sub> °C	741.4	757.8	737.9	747.7	717.1	737.4	739.88		
P kbar	3.81	3.54	4.11	3.46	2.77	2.60	3.38		
Depth km	10.27	9.57	11.08	9.33	7.48	7.01	9.12		
r1	-3.80	-4.18	-2.69	-3.07	-2.49	-2.87	-3.18		
r2	-6.69	-6.57	-5.16	-5.64	-5.01	-4.78	-5.64		

US AFG: Umm Shommer Alkali-feldspar granites; SH AFG: Sharm Alkali-feldspar granite (from Eyal et al., 2010); AI: Apatitic Index = molar Al/(Na+K); T<sub>Zr</sub> °C: Zircon saturation thermometer calculated according to Watson & Harrison 1983 and Boehnke et al. 2013; P kbar and Depth km: Crystallization pressures and depth of emplacement were calculated according to Yang (2017); r1, r2 discrimination values calculated by Agrawal (1995).

### Conditions of crystallization

The crystallization temperatures of HGAFGs were calculated using zircon saturation (T<sub>Zr</sub> °C) thermometer (Watson & Harrison, 1983; Boehnke et al., 2013). They range from 717 to 758 °C. The pressures of crystallization of HGAFGs were estimated by using the following equation proposed by Yang (2017):

$$P = 0.2426 \times (ab + or)^3 - 46.397 \times (ab + or)^2 + 2981.3 \times (ab + or) - 64224$$

The pressures range from 2.60 to 4.11 (Table 2), while the depths of emplacement of HGAFGs were calculated depending on the estimated pressure by using Yang (op. cit.) calculations of the depth, which suppose that 1 kb, equivalent to 2.7 km depth assuming that, the density of continental crust is 2.85 g/cm<sup>3</sup> and the pressure is entirely lithostatic. They range from 7 to 11 km (Table 2).

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Fig. 4: (a-e) a)  $10^*TiO_2 - FeO + MnO - MgO$  diagram (Nachit et al., 2005), distinguishing between magmatic, re-equilibrated and neofomed biotite; b)  $Al^{IV}$  versus  $Fe^{2+}/(Fe^{2+}+Mg)$  diagram (Rieder et al., 1999) for classification of the studied trioctahedral micas; c)  $Al^{IV}$ - $Mg$  biotite diagram (Nachit et al., 1985); d)  $Fe^{2+}-Fe^{3+}-Mg$  diagram (Wones & Eugster, 1965), Oxygen fugacity buffers: quartz - fayalite- magnetite (QFM), nickel - nickel oxide (NNO), hematite-magnetite (HM); e)  $\log fO_2 - T^{\circ}C$  ternary diagram for the biotite stability as a function of  $Fe/(Fe+Mg)$  at pressure of 2070 bars (Wones & Eugster, 1965), Arrows represent oxidizing and reducing trends.

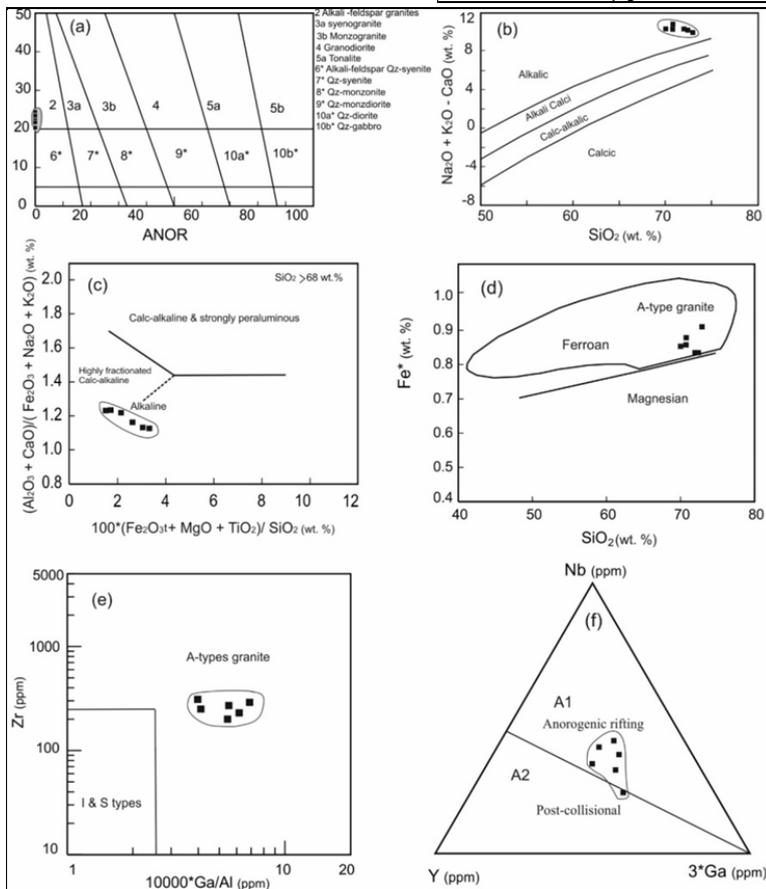
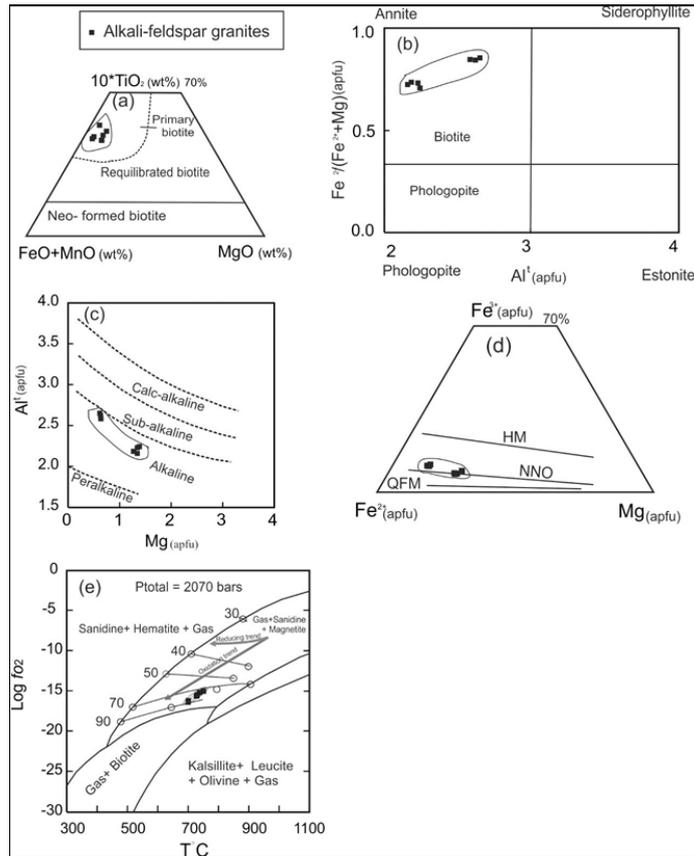


Fig. 5. a) ANOR-Q diagram (Streckeisen & Le Maitre, 1979); b)  $SiO_2 - Na_2O + K_2O - CaO$  diagram (Frost & Frost, 2008; Frost et al., 2001); c) major elements discrimination diagram (Sylvester, 1989); d)  $SiO_2 - Fe^*$  diagram (Frost et al., 2001)  $Fe^*$ :  $Fe \text{ number} = (Fe_2O_3 t / Fe_2O_3 t + MgO)$ ; e)  $10000^*Ga/Al$ -Agpaitic Index (AI) diagram (Whalen et al., 1987); f)  $Y-Nb-3^*Ga$  ternary diagram (Eby, 1992).

**DISCUSSION**

**Tectonic setting**

On r1-r2 multivariate discriminant diagram (Agrawal, 1995), HGAFGs fall in anorogenic granite field (Fig. 6a). In SiO<sub>2</sub>-log [CaO/(Na<sub>2</sub>O + K<sub>2</sub>O)] diagram (Brown, 1982), HGAFGs samples are plotted on extensional tectonic setting (Fig. 6b). According to K<sub>2</sub>O-Na<sub>2</sub>O- 3\* CaO ternary diagram (Liégeois et al., 1998). HGAFGs fall in Late-Shear Alkaline granite series (Fig. 6d). Alkaline magmatism is linked to major lithospheric structures, whether it is post-collisional (Liégeois & Black, 1987; Duchesne et al., 1997) or anorogenic (Black et al., 1985; Moreau et al., 1994). Alkaline magmatism is not always present in post-collisional settings, but when it occurs it is very similar to true anorogenic alkaline manifestations (Black et al., op. cit.; Sylvester, 1989). Nb/Y- Rb/Y diagram (Pearce et al., 1990), the HGAFGs samples plot in between lower crust composition and they follow Within-Plate enrichment trend (Fig. 6c). According to Sr-Rb diagram (Condie, 1973), HGAFGs magmatic source was generated at depth more than 30 km (Fig. 6d) and emplaced later at shallower depth. The ANS crustal thickness is around 40 km (Al-Damegh et al., 2005).

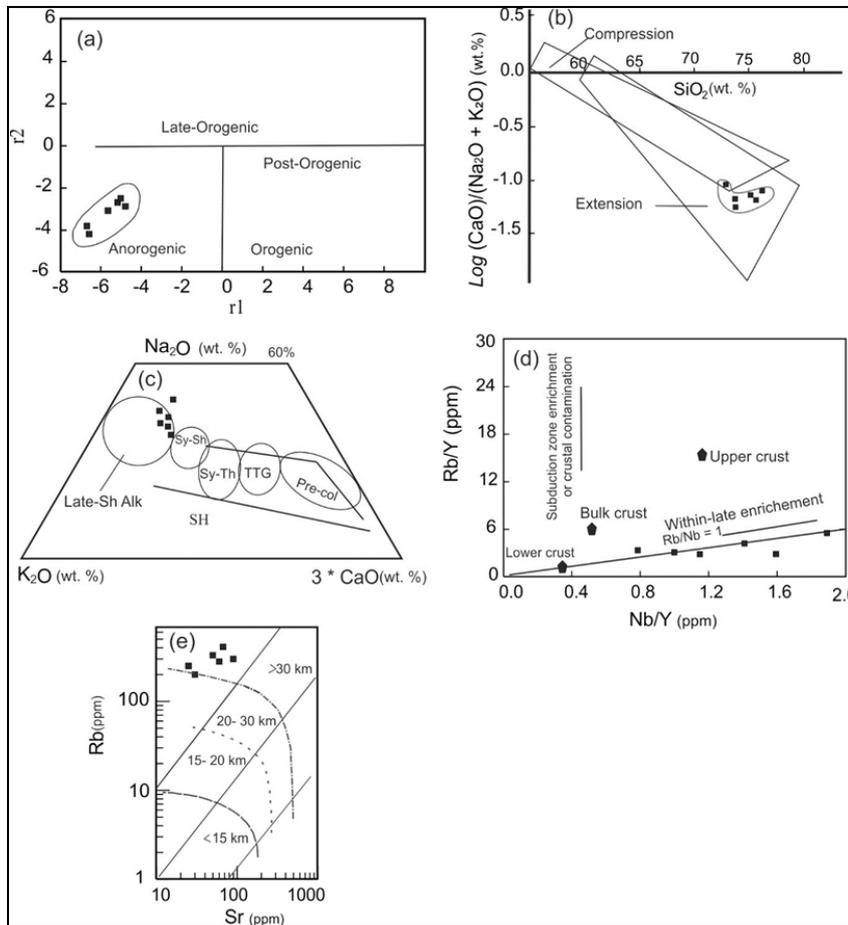


Fig. 6. a) r1-r2 multivariate discriminant diagram (Agrawal, 1995); b) SiO<sub>2</sub>-log [CaO/( Na<sub>2</sub>O + K<sub>2</sub>O)] diagram (Brown, 1982); c) K<sub>2</sub>O-Na<sub>2</sub>O- 3\* CaO ternary diagram (Liégeois et al., 1998); d) Nb/Y- Rb/Y diagram, the compositions of lower-, bulk-, and upper crusts are after Taylor & McLennan (1985); e) Sr-Rb binary diagram fields after Condie (1973).

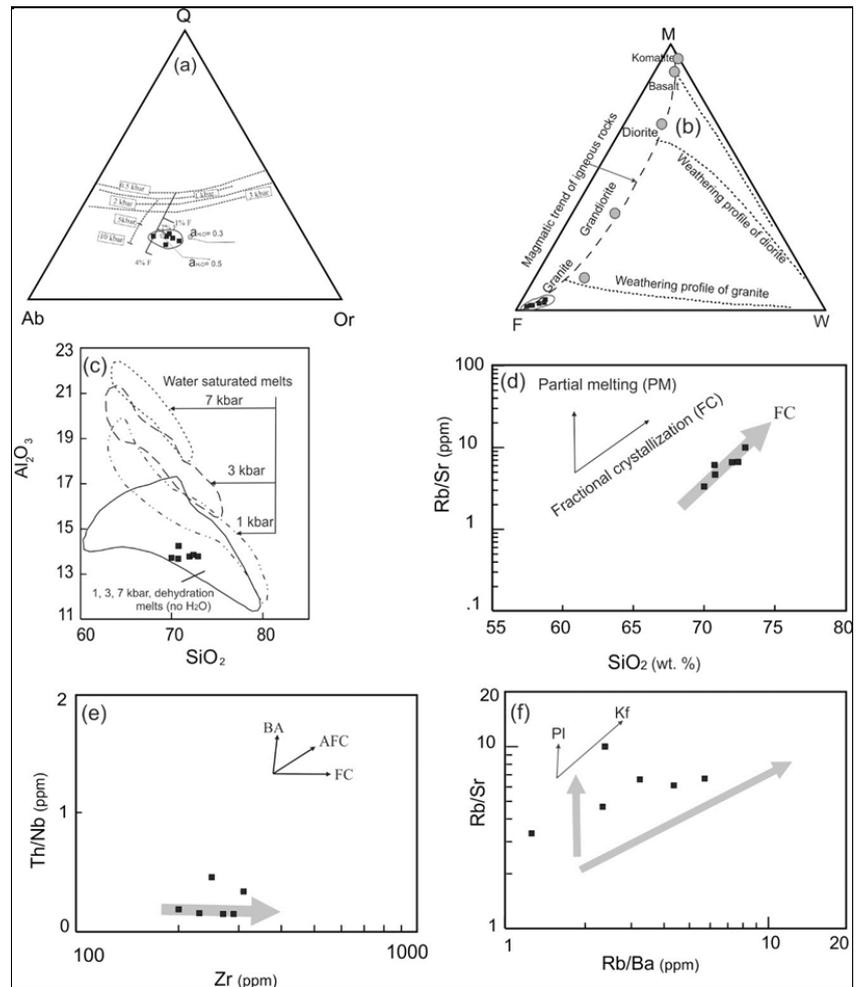
**Petrogenesis**

Ab- Q- Or-H<sub>2</sub>O- F plots of normative compositions show that the HGAFGs fall just close to the minima melt at 5 kbar with water pressure a<sub>H<sub>2</sub>O</sub> 0.5 and at 2%, F (Fig. 7a), suggesting that the HGAFGs does not represent 100% liquid composition, or that the melt was water-undersaturated (Johannes & Holtz, 1990). This trend is consistent with the moving positions of the minima for the water-saturated Q-Ab-Or system with increasing fluorine and water content. Many studies refer to the role halogens (F and Cl) as complexing agents for REE, Y and HFSE (Charoy & Raimbault, 1994; Audétat et al., 2000; Agangi et al., 2010). The origin of the F-rich fluids either are primary magmatic origin (Dostal & Chatterjee, 1995;

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Agangi *op. cit.*) or as secondary metasomatic origin (Nurmi & Haapala, 1986). HGAFGs are fluorine-rich (1.5 F wt. %, on average; Table 2). The appearance of fluorite in CIPW normative values (Table 2) and presence fluorite, support that, the fluorine has a role of in the genesis of these granites. In addition, MFW plot of Ohta and Arai (2007) used to differentiate between fresh magmatic and altered samples. All HGAFGs samples follow the magmatic trend (Fig. 7b). In addition, there is no any unusual F-rich rocks surrounds HGAFGs to be a source for leaching fluorine through hydrothermal processes. Furthermore, the lack of any metasomatic alteration (i.e. albitization) support the magmatic origin of F-rich fluids. Agangi *et al.* (2010) supposed the following scenario of F-rich fluids formation, which start with the crystallization of quartz and feldspar from the magmatic source, yielding a late-stage magmatic fluid enriched in F and to lesser extent in water content with fractionation of anhydrous quartzo-feldspathic components. Fluorine dissolved in the magma during late stages of magmatic evolution, causing REE and HFSE complexing and mobilizing.

Fig. 7. a) Ab-Q-Or-H<sub>2</sub>O-F, dotted lines show the location of minima melt composition at saturated water pressure ranging from 0.5 to 10 kbar (Winkler *et al.*, 1975). Solid line shows the minima melt at 1 kbar with excess H<sub>2</sub>O at 1%, 2%, 4% F (Manning, 1981); b) FMW diagram (Ohta and Arai, 2007) for chemical weathering of igneous rock M: mafic, F: felsic W: weathering; c) SiO<sub>2</sub>-Al<sub>2</sub>O<sub>3</sub> diagram represents comparison between the compositions of the investigated granites with the experimental work of the melting of a basaltic/amphibolitic source (Bread and Lofgren, 1991); d) SiO<sub>2</sub>-Rb/Sr binary diagram (Blevin 2004), PM: partial melting, FC: fractional crystallization trends from (Pearce *et al.*, 1990); e) Zr-Th/Nb diagram, trends of FC: fractional crystallization, AFC: assimilation-fractional crystallization, and BA: bulk assimilation are from Nicolae and Saccani (2003); f) Rb/Ba-Rb/Sr diagram, fractionation trends of Pl: plagioclase, Kf: k-feldspar.



Finally, these F-enriched fluids migrate toward the roof of the pluton, giving rise to formation of accessory minerals, in interstices spaces between early formed minerals (i.e. feldspar and quartz). According to Moreno *et al.* (2014) the F-rich melt could be formed through carbonatite and amphibole-rich mantle wedge as it broke down generates F-rich melt.

SiO<sub>2</sub>-Al<sub>2</sub>O<sub>3</sub> binary diagram (Bread & Lofgren, 1991) shows the compositional similarities between the investigated HGAFGs analyses and experimental results of 20-25 % dehydration melting of amphibolitic/basaltic sources (i.e. no H<sub>2</sub>O added). HGAFGs samples fall in the field of amphibolitic source at 1, 3, 7 kbar (Fig. 6c). Both SiO<sub>2</sub>- Rb/Sr (Fig. 7d after Blevin, 2004) and Zr- Th/Nb (Fig. 6e after Nicolae and Saccani, 2003) diagrams reveal that, fractional crystallization was probably the main process

to be accounted for the evolution of HGAFGs. Fractionation of plagioclase and K-feldspar could be responsible for the depletion of Ba (from 240 to 35 ppm) and Sr (90 to 25 ppm) and the enrichment of Rb (from 200 to 410 ppm, cf. Table 2 and Fig. 7f).

### Geodynamic implications

The widely distributed calc-alkaline and alkaline coeval rocks in the ANS explained by extensional collapse that follows continental collision (650–600 Ma) between the East and West Gondwana (Abdelsalam et al. 2002, Fig. 8a). From one side, the continental collision led to extensive crustal thickening (Stern 1994, 2002; Abdelsalam & Stern 1996; Stoesser & Frost 2006; Stern 2008; Avigad & Gvirtzman 2009, Fig. 8b). On the other side, the extensional collapse (600-550Ma) led to thinning of lithosphere (Fig. 7c). Extensional collapse was controlled by lithospheric delamination and slab breakoff (Davies & von Blackenburg, 1995; Farahat et al. 2007; Avigad & Gvirtzman op. cit.; Eyal et al. 2010).

During late collisional stage (630-610 Ma), due to the beginning of the slab breakoff, a slab window opened, which enhancing the heat flux provided by the upwelling asthenosphere, causing melting process (Bonin, 2004; Fig. 8b). This melts produced late-collisional calc-alkaline magmatism in the northern ANS (Fig. 8b).

The next late collisional to early post-collisional stage (610-590 Ma) commenced ~20 m.y. later on, the rapid uplift of the crust due to the slab detachment caused decompression melting in the lithosphere. This mafic mantle-derived melt underplate the juvenile crust, causing widespread magmatism and producing coeval calc-alkaline/alkaline magmas (Fig. 8c).

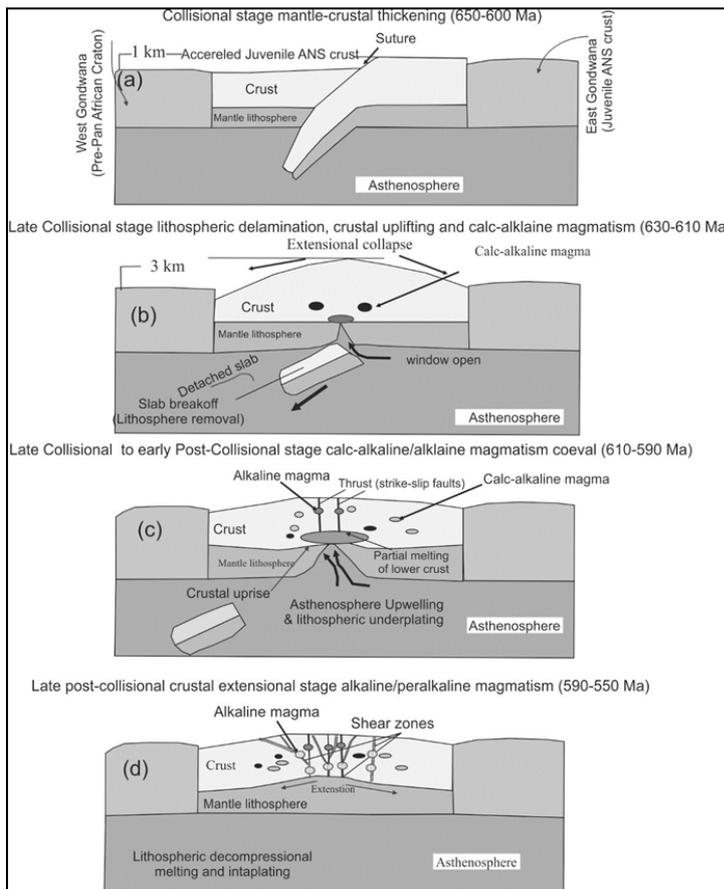


Fig. 8: Sketch showing the tectono-magmatic evolution of late post-collisional alkali-feldspar granites of Homrat El-Girigab (modified after Elwan et al., 2019), a) Collision between East and West of Gondwana (650–600 Ma), which characterized by mantle/crust thickening; b) Beginning extensional collapse at ~ 600 Ma and slab breakoff as a consequence of crustal thickening. Due to slab breakoff a slab window opened, which enhance heat flux provided by the upwelling asthenosphere, causing melting. This melts produced late-collisional calc-alkaline magmatism; c) Slab detaches and sink away and the detached slab or lithospheric removal allows asthenosphere to rise or asthenosphere upwelling causing melting and formation of calc-alkaline /alkaline magmatism; d) Rapid lithospheric uplift and intraplate caused melting to produce alkaline/peralkaline magmatism in ANS, which facilitated by strike-slip faults and shear-zones.

During the late post-collisional crustal extensional stage (590-550 Ma), the effect of lithospheric delamination, and thus asthenospheric uprise, likely diminishes. Due to the rapid uplift mantle-derived

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melts intraplate the middle-lower crustal levels forming alkaline/peralkaline magmatism of HGAFGs, which facilitated by the abundance of strike-slip faults and shear zones (Fig. 8d)

As the HGAFGs supposed to represent A<sub>1</sub> anorogenic granites, they most probably originated in crustal extension regime (Grebennikov, 2014). They formed in a within-plate settings; either in the intracontinental system or near the divergent boundaries, where formation of igneous rocks is related to faulting, shear zones and mantle melts penetration (Dobretsov, 2003). HGAFGs represent late shear alkaline granites (Figs. 5c, and 7d), where mantle-derived mafic melts intraplate the lower crustal levels, facilitated by the abundance of strike-slip faults and shear zones. HGAFGs were generated from lower crustal amphibolitic source through fractional crystallization (Figs. 5 d & e, 6 c, d & f). The abundant amphibole and carbonatite in the mantle wedge (Moreno et al., 2014) broke down to generate F-rich melt. Partial melting of the hot juvenile lower crust injected with such melt then produced granite with A-type characteristics. The mantle components in the mafic melt were metasomatized by subduction released fluids (i.e. F-rich melts) are easier to be partially melted, which consequently eliminated subduction signatures and formed HGAFGs magma.

### CONCLUSION

- 1- Homrat El-Girigab area located at northern Eastern Desert and it is covered by Dokhan volcanics, which intruded by alkali-feldspar granites.
- 2- The biotite in Homrat El-Girigab alkali-feldspar granites were crystallized from alkaline source under NNO Oxygen buffer and oxygen fugacity (log fO<sub>2</sub>) range from -14.92 to -16.28 i.e. oxidized conditions.
- 3- Homrat El-Girigab alkali-feldspar granites are alkaline, ferroan anorogenic A-type granites. They are belong to A1-subtype anorogenic granites.
- 4- Alkali-feldspar granites were crystallized under temperatures range from 700 to 758°C, pressures 3 to 4 kbar and emplaced at depths range from 7 to 11 km.

The alkali-feldspar granites were emplaced during the late post-collisional crustal extensional stage, where mantle-derived mafic melts intraplate the lower crustal levels, which facilitated by the abundance of strike-slip faults and shear zones. Homrat El-Girigab granites were derived from lower crustal amphibolitic source and evolved by fractional crystallization.

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## نشأة جرانيتات الفلسبار القلي بمنطقة حمرة الجرجاب - شمالي الصحراء الشرقية - مصر

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تقع منطقة حمرة الجرجاب في شمالي الصحراء الشرقية والذي يتميز بانتشار المتداخلات الجرانيتية الكالس- قليه وقلية فوق القليه بالإضافة الى مصاحباتها من الصخور البركانية. هذه الجرانيت لها أهمية جيوديناميكية خاصة حيث انها ممكن ان تمدنا بمعرفة عن كيفية تكون القشرة القارية للدرع العربي النوبي . منطقة حمرة الجرجاب مغطاة ببركانيات الدخان (انديزيت و داسيت) والتي قطعت بجرانيتات الفلسبار القلي .

تشير كيميائية معدن البيوتيت الى ان جرانيتات الفلسبار القلي قد تبلرت من ماجما قليه وذلك تحت ظروف أكسدة وقد تبلرت جرانيتات الفلسبار القلي في ظروف من درجات الحرارة التي تراوحت بين ٧٠٠-٧٥٠ درجة مئوية، والضغط تراوحت بين ٣-٤ كيلوبار وتداخلت على عمق يتراوح بين ٧ الى ١١ كيلومتر وتحت لوغاريتم نشاط الأكسجين يتراوح بين ١٥- إلى ١٦ .

جرانيتات الفلسبار القلي بحمرة الجرجاب هي صخور قليه غنية بالحديد وقد تكونت في بيئه لا اوروجينية أو تباعديه من نوع (A-type) وذلك خلال فترة متأخرة من مرحلة ما بعد التصادم حيثما بدأ يضمحل تأثير كل من فصل صفيحة الليثوسفير و بروز الاستينوسفير ، في هذا المرحلة بدأت الماجما المستمدة من الستارة في التخلل في الجزء السفلى من القشرة الأرضية ، وهذا التخلل اصبح اكثر سهولة من خلال وفرة كل من صدوع الدسر ونطاقات القص. وقد سبب هذا التخلل انصهار واسع منتجا الصهير القلي التي كونت جرانيتات حمرة الجرجاب. وقد نشأت الجرانيتات تحت دراسته من مصدر امفيبوليتي التركيب في الجزء السفلى من القشرة وقد تطورت هذه الجرانيتات غالبا عن طريق التبلر التجزيئي.